High energy radiation profile of Jupiter as observed by the JunoASC

Jørgensen, Finn E.; Jørgensen, J. L.; Jørgensen, P. S.; Herceg, M.; Denver, T.; Benn, M.; Shushkova, J.; Connerney, J. E. P.; Oliversen, R. J.; Gershman, D. J.

Publication date:
2018

Document Version
Publisher's PDF, also known as Version of record

Link back to DTU Orbit

Citation (APA):
High energy radiation profile of Jupiter
as observed by the JunoASC

Finn E. Jørgensen¹, J.L. Jørgensen¹, P.S. Jørgensen¹, M. Herceg¹, T. Denver¹, M. Benn¹, J. Shushkova¹, J.E.P. Connerney²,³, R.J. Oliversen², D. J. Gershman², S. Kotsiaros²

¹Technical University of Denmark (Denmark). ²NASA Goddard Space Flight Center (USA). ³Space Research Corporation (USA)

The Magnetometer Investigation’s star trackers (ASCs) on Juno spacecraft are collocated with the magnetometer sensors on a boom. Magnetometer optical bench allowed for limited radiation shielding of the cameras. In addition to suppressing the undesirable effects of penetrating energetic particles, the morphological filters in the instrument flight software provide a measurement of the actual penetrating flux at all times.

In addition to suppressing the undesirable effects of The ASC has proﬁled three distinct regions of high energy particle ﬂuxes: the main plasma torus, the plasma populating the magnetic ﬁeld lines reaching the auroral region, and, particle ﬂuxes on near open ﬁeld lines. We present a map of the regions proﬁled this far, and their associated ﬂuxes, we show the high energy ﬂuxes observed on ﬁeld lines connecting to the Jovian satellites, and we discuss the particle nature observed on open ﬁeld lines.

**Micro Advanced Stellar Compass µASC**
- Designed and produced by the Measurement and Instrumentation Systems (GSIU)
- One of the most successful star tracker worldwide
- Autonomously calculates altitude based on all bright stars in the CHUs Field of View
- Provides 22 true solutions per second with one single CHU
- Absolute accuracy of < 1 arc second
- Operating on many satellite missions without a single hardware or functional failure

**Calibration**
The µASC response to the high energy irradiation is fully calibrated using 1-2 MeV electron and 30-200 MeV proton irradiation with varied incidence angle, flux, dose rate, temperature, TID, and DDD doses.

**Image analysis. Deposited energy**
A track left by an energetic particle that hit the CCD is ~50 pixels long (~0.7 mm or 160 MeV/cm²) and has 4 or 5 splits due to recoil nuclei produced by the incident particle in silicon. Estimation of the primary energy of a particle that could traverse at least 0.7 mm in Si behind the heavy shielding will involve the transport flux modelling for the µASC, the calibration curves of the CCD and inﬂight image analysis.

**Observations**
The wiggle plot below shows Juno’s position in magnetic coordinates (Rho, Z) in [R̃₁]. In the trace of the orbit, the flux measured by the µASC is displayed with a color scale. The major Jovian moons’ radial distance have been also added in the plot in addition to the model of the magnetodisc.

Juno’s crossing of the magnetic field lines which intersect the Jovian moon orbits is shown. These crossings are correlated with observations of the µASC particle counter.

**Jovian magnetic field**
Jupiter’s magnetic field model (JRM09) by Connerney (2018) describes the ﬁeld with spherical harmonics up to degree 10. It is based on measurements performed near Jupiter by the Juno mission. The JRM09 model shows a more complex Jovian magnetic ﬁeld in comparison to earlier models (eg. VPI4), and a clear asymmetry within the northern and southern hemispheres. For distances to the planet (~10R̃), the magnetic ﬁeld can be represented well by a 10.5” tilted magnetic dipole and there is no apparent differences wrt VPI4 model.

**Highlights**
- Model of the high energy electron radiation environment near the
- Better constrains on the Jupiter’s current sheet
- Magnetic footprint of the Jupiter’s inner planet


American Geophysical Union, Fall meeting, Washington, D.C., 10-14 Dec 2018

fj@space.dtu.dk